

[1511.05220], [1511.05975] with Seungho Gwak, Karapet Mkrtchyan, SooJong Rey

RAINBOW Valley
of
Colored (HS) Gravity

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100 years old

Defy Einstein Gravity?



- Super Gravity
- Higher-derivative Gravity
- Massive Gravity
- **Higher Spin Gravity**
- **Colored Gravity**



Einstein Gravity \longleftrightarrow **a single** massless spin 2

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Colored Gravity:

Multiple massless spin 2
with color decoration



Einstein Gravity \longleftrightarrow **a single** massless spin 2

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Q1 Isn't it wrong? No, it works in certain cases

Q2 Isn't it straightforward and boring?

There are **surprising** features

No-Go for colored gravity



- Multiple massless spin 2 $h_{\mu\nu}^I$

👉 cubic interactions

$$g_{IJK} \left(h_{\mu\rho}^I \partial^\rho h_{\nu\lambda}^J \partial^\lambda h^{K\mu\nu} + \dots \right)$$

symmetric

👉 global symmetries

$$[M_{\mu\nu}^I, M_{\rho\lambda}^J] = 4 g^{IJ} g_{K\lambda} \eta_{[\nu[\rho} M_{\lambda]\mu}^K]$$

associative

⇒ **only trivial solution**

No-Go for colored gravity



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- Color-charged massless spin 2 in flat space

👉 minimal interaction to gauge field

⇒ **violation of color gauge inv.**

No-Go for colored gravity



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- Color-charged massless spin 2 in flat space **(AdS)**

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Yes-Go for colored gravity

Global Symmetries

isometry $\mathfrak{g}_i \otimes \mathfrak{g}_c$ color sym.
 M_X T_I



Yes-Go for colored gravity

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$$[M_X \otimes T_I, M_Y \otimes T_J] = \frac{1}{2} [M_X, M_Y] \otimes \{T_I, T_J\} + \frac{1}{2} \{M_X, M_Y\} \otimes [T_I, T_J]$$

not defined for Lie algebra



Yes-Go for colored gravity

Global Symmetries

$$\begin{array}{ccc} \textit{Associative} & & \textit{Associative} \\ \text{isometry} & \mathfrak{g}_i \otimes \mathfrak{g}_c & \text{color sym.} \\ M_X & & T_I \end{array}$$

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- \mathfrak{g}_c associative color algebra: $\mathfrak{u}(N)$
- \mathfrak{g}_i associative algebra containing isometry



Yes-Go for colored gravity

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- \mathfrak{g}_c associative color algebra: $\mathfrak{u}(N)$
 - \mathfrak{g}_i associative algebra containing isometry
- ➡ contains more spectrum (e.g. HS)

3D Chern-Simons

Colored Gravity

- CS action: $S[\mathcal{A}] = \frac{\kappa}{4\pi} \int \text{Tr} \left(\mathcal{A} \wedge d\mathcal{A} + \frac{2}{3} \mathcal{A} \wedge \mathcal{A} \wedge \mathcal{A} \right)$
- Gauge Algebra: $\mathfrak{g} = (\mathfrak{gl}_2 \oplus \mathfrak{gl}_2) \otimes \mathfrak{u}(N) \oplus \text{id} \otimes \mathbf{I}$

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two additional gauge fields

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subtract Abelian CS
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subtract Abelian CS

two additional gauge fields

$$[M_{ab}, M_{cd}] = 2 (\eta_{d[a} M_{b]c} + \eta_{c[b} M_{a]d}), \quad [M_{ab}, P_c] = 2 \eta_{c[b} P_{a]}, \quad [P_a, P_b] = \sigma M_{ab}$$

$$M_{ab} = \frac{1}{2} \epsilon_{ab}{}^c (J_c + \tilde{J}_c), \quad P_a = \frac{1}{2\sqrt{\sigma}} (J_a - \tilde{J}_a)$$

$$\begin{aligned} \sigma &= +1 \text{ for AdS}_3 \\ \sigma &= -1 \text{ for dS}_3 \end{aligned}$$

$$\text{Tr}(J_a J_b) = 2\sqrt{\sigma} \eta_{ab}, \quad \text{Tr}(\tilde{J}_a \tilde{J}_b) = -2\sqrt{\sigma} \eta_{ab}, \quad \text{Tr}(\mathbf{T}_I \mathbf{T}_J) = \delta_{IJ}$$

3D Chern-Simons

Colored Gravity

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☹ not tangible

☹ is it over?

☹ so what?



3D Chern-Simons

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Let's rewrite this in **METRIC** form!

solve torsion condition only for the genuine graviton (singlet spin two)

3D Colored Gravity

 **solve torsion condition** 

3D Colored Gravity

👉 solve torsion condition 👈

$$S = S_{\text{CS}} + \frac{1}{16\pi G} \int d^3x \sqrt{|g|} \left[R - V(\varphi, \tilde{\varphi}) + \frac{2\sqrt{\sigma}}{N\ell} \epsilon^{\mu\nu\rho} \text{Tr} \left(\varphi_{\mu}{}^{\lambda} D_{\nu} \varphi_{\rho\lambda} - \tilde{\varphi}_{\mu}{}^{\lambda} D_{\nu} \tilde{\varphi}_{\rho\lambda} \right) \right]$$

3D Colored Gravity

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- SU(N) CS: $S_{\text{CS}} = \frac{\kappa\sqrt{\sigma}}{2\pi} \int \left[\text{Tr} \left(\mathbf{A} \wedge d\mathbf{A} + \frac{3}{2} \mathbf{A} \wedge \mathbf{A} \wedge \mathbf{A} \right) - \text{Tr} \left(\tilde{\mathbf{A}} \wedge d\tilde{\mathbf{A}} + \frac{3}{2} \tilde{\mathbf{A}} \wedge \tilde{\mathbf{A}} \wedge \tilde{\mathbf{A}} \right) \right]$
must be quantized!

3D Colored Gravity

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- Newton's constant: $\kappa = \frac{\ell}{4NG}$

semi-classical gravity: compatible with small CS level for large N!

3D Colored Gravity

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- Colored spinning matter: $D_{\mu} \varphi_{\nu\rho} = \nabla_{\mu} \varphi_{\nu\rho} + [A_{\mu}, \varphi_{\nu\rho}]$
covariant couplings!

Potential

$$V(\varphi, \tilde{\varphi}) = -\frac{1}{N\ell^2} \text{Tr} \left[2\sigma \mathbf{I} + 4 \left(\varphi_{[\mu}^{\mu} \varphi_{\nu]}^{\nu} + \tilde{\varphi}_{[\mu}^{\mu} \tilde{\varphi}_{\nu]}^{\nu} \right) + 8\sqrt{\sigma} \left(\varphi_{[\mu}^{\mu} \varphi_{\nu}^{\nu} \varphi_{\rho]}^{\rho} - \tilde{\varphi}_{[\mu}^{\mu} \tilde{\varphi}_{\nu}^{\nu} \tilde{\varphi}_{\rho]}^{\rho} \right) \right] \\ - \frac{16\sigma}{N^2\ell^2} \text{Tr} \left(\varphi_{[\mu}^{\nu} \varphi_{\rho]}^{\rho} - \tilde{\varphi}_{[\mu}^{\nu} \tilde{\varphi}_{\rho]}^{\rho} \right) \text{Tr} \left(\varphi_{[\nu}^{\mu} \varphi_{\lambda]}^{\lambda} - \tilde{\varphi}_{[\nu}^{\mu} \tilde{\varphi}_{\lambda]}^{\lambda} \right) + \frac{6\sigma}{N^2\ell^2} \left[\text{Tr} \left(\varphi_{[\mu}^{\mu} \varphi_{\nu]}^{\nu} - \tilde{\varphi}_{[\mu}^{\mu} \tilde{\varphi}_{\nu]}^{\nu} \right) \right]^2$$

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Matter self-interaction: \sqrt{N} times **STRONGER** than gravity!

Closer look on Potential

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- reduction to parity-invariant sector

$$\chi_{\mu\nu} = \sqrt{\sigma} (\varphi_{\mu\nu} - \tilde{\varphi}_{\mu\nu}) \\ \tau_{\mu\nu} = \varphi_{\mu\nu} + \tilde{\varphi}_{\mu\nu}$$

$$V(\chi) = -\frac{2 \sigma}{N \ell^2} \text{Tr} \left(\mathbf{I} + \chi_{[\mu}{}^\mu \chi_{\nu]}{}^\nu + \chi_{[\mu}{}^\mu \chi_{\nu}{}^\nu \chi_{\rho]}{}^\rho \right)$$

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- reduction to diffeo-invariant sector

$$\chi_{\mu\nu} = g_{\mu\nu} \mathbf{X}$$

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Closer look on Potential

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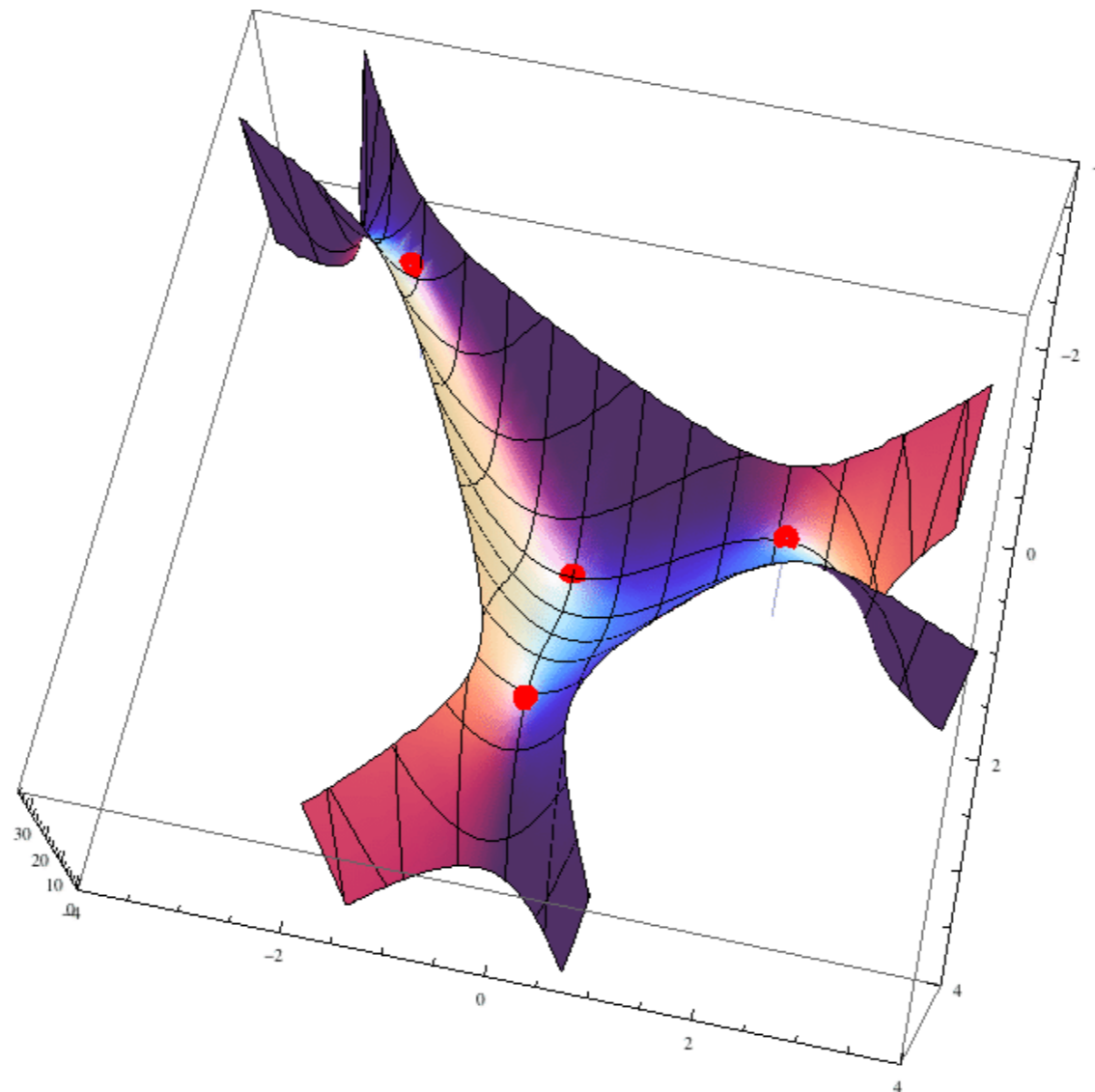
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Non-trivial Potential with Many Extrema!

Rainbow Vacua

$$V(\mathbf{X}) = -\frac{2\sigma}{N\ell^2} \text{Tr}(\mathbf{I} + 3\mathbf{X}^2 + \mathbf{X}^3)$$

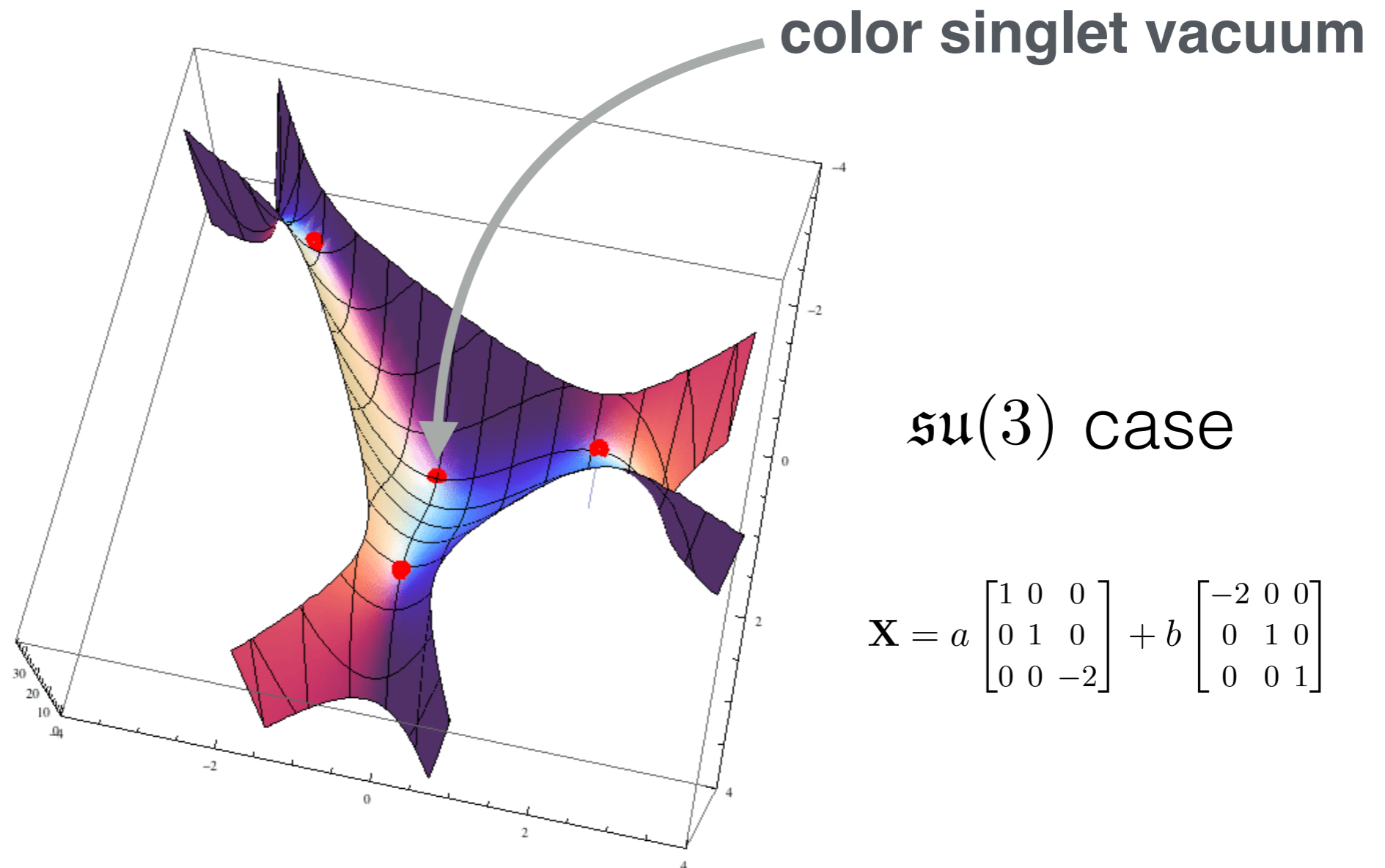


$\mathfrak{su}(3)$ case

$$\mathbf{X} = a \begin{bmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & -2 \end{bmatrix} + b \begin{bmatrix} -2 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

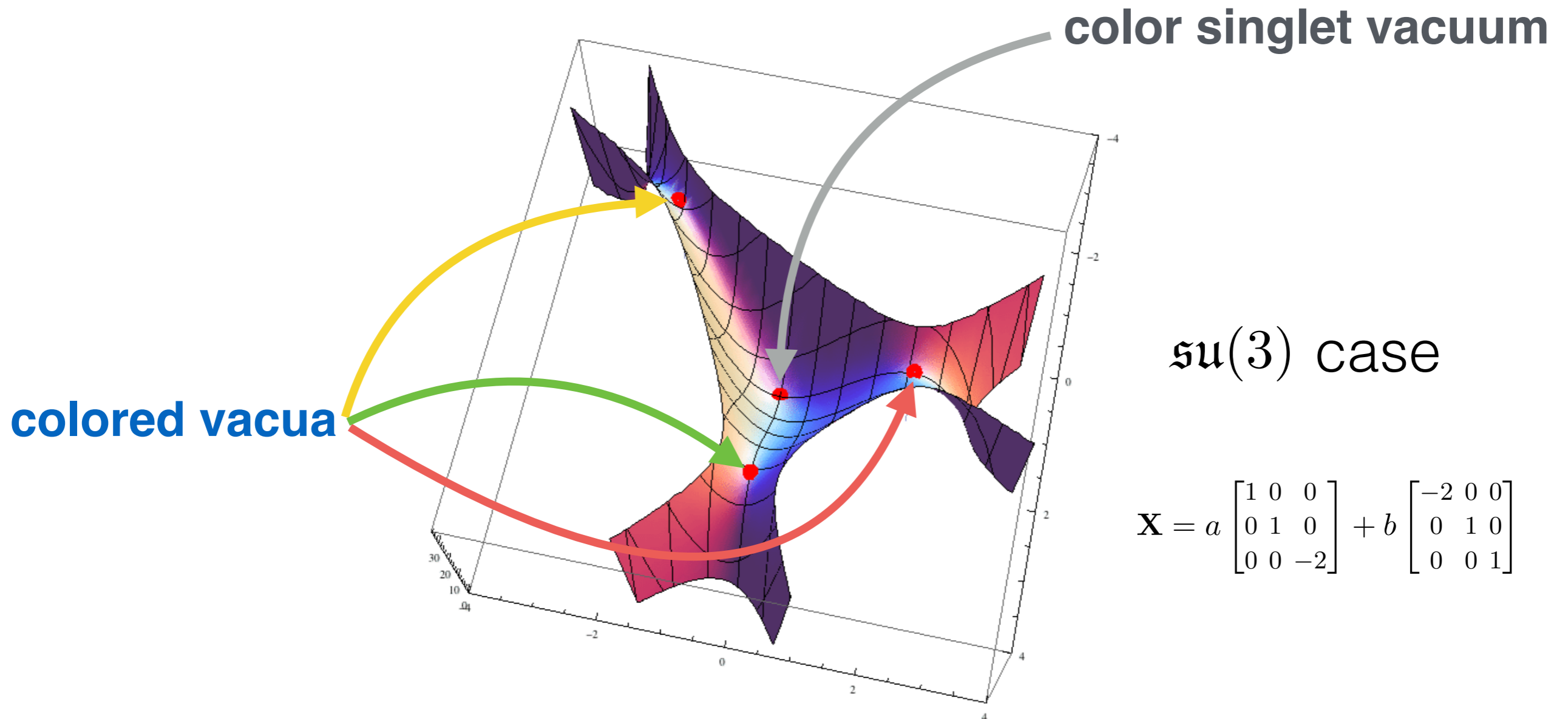
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- solutions up to SU(N) rotation:

$$\mathbf{X} = \frac{N}{\text{Tr}(\mathbf{Z})} \mathbf{Z} - \mathbf{I} \quad \mathbf{Z}_k = \begin{bmatrix} \mathbf{I}_{(N-k) \times (N-k)} & 0 \\ 0 & -\mathbf{I}_{k \times k} \end{bmatrix}$$

Parameter

$$k = 0, 1, \dots, \left[\frac{N-1}{2} \right]$$

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Symmetry Breaking

$$SU(N - k) \times SU(k) \times U(1)$$

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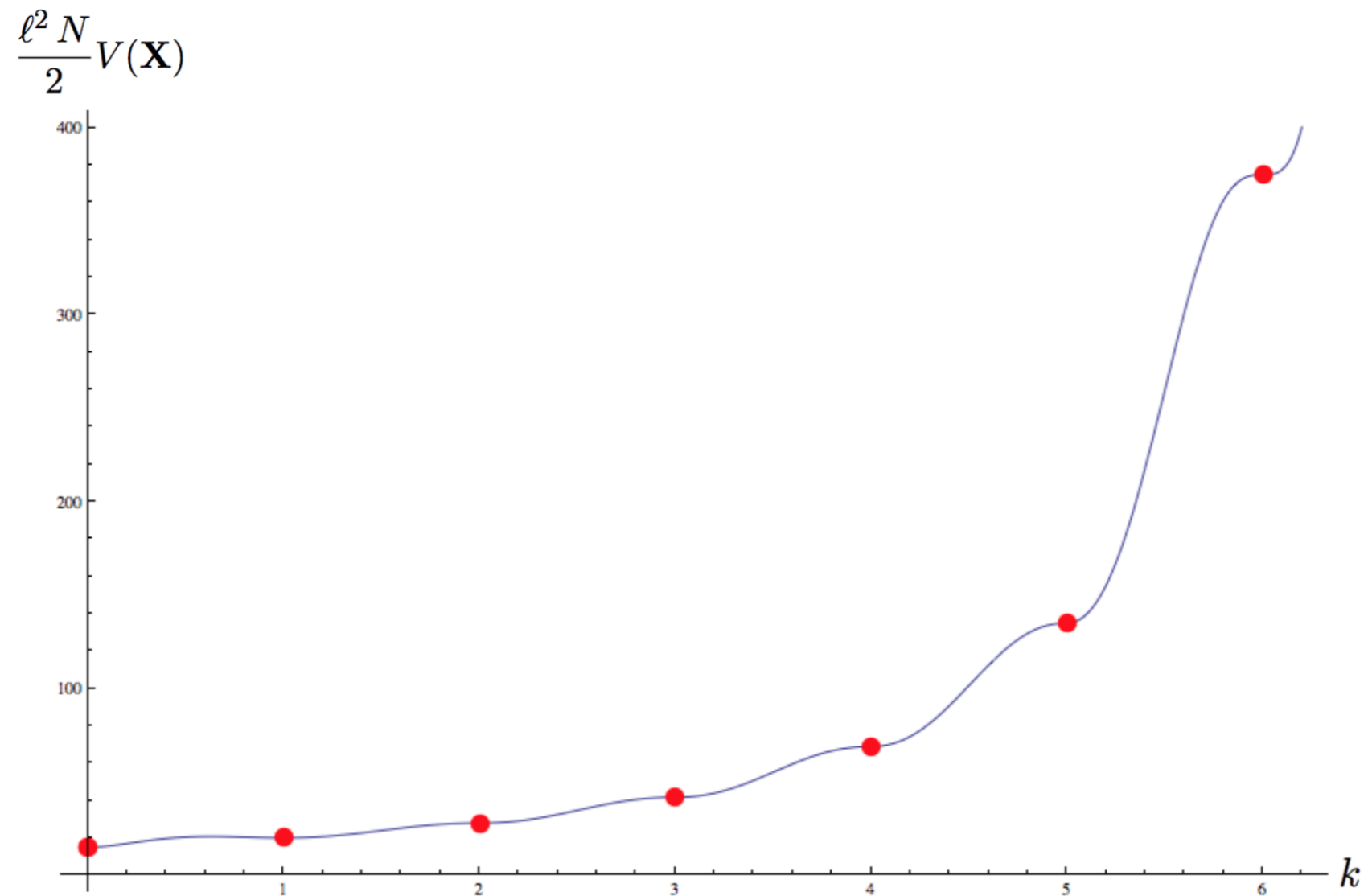
$$SU(N-k) \times SU(k) \times U(1)$$

- extremum values: $V(\mathbf{X}_k) = -\frac{2\sigma}{\ell^2} \left(\frac{N}{N-2k} \right)^2 = 2\Lambda_k$

Vacuum dependent Cosmological Constant!

Rainbow Vacua

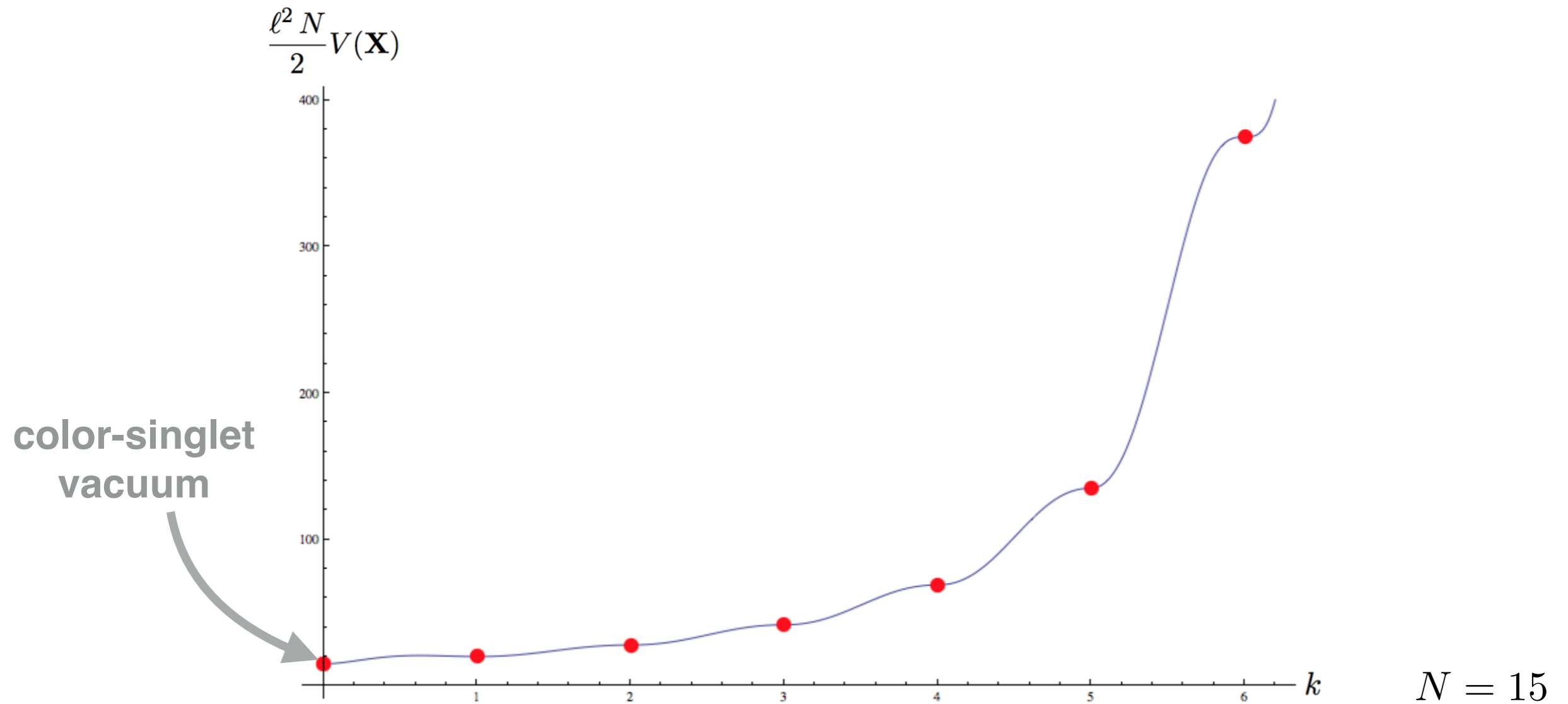
$$V(\mathbf{X}) = -\frac{2\sigma}{N\ell^2} \text{Tr}(\mathbf{I} + 3\mathbf{X}^2 + \mathbf{X}^3)$$



$N = 15$

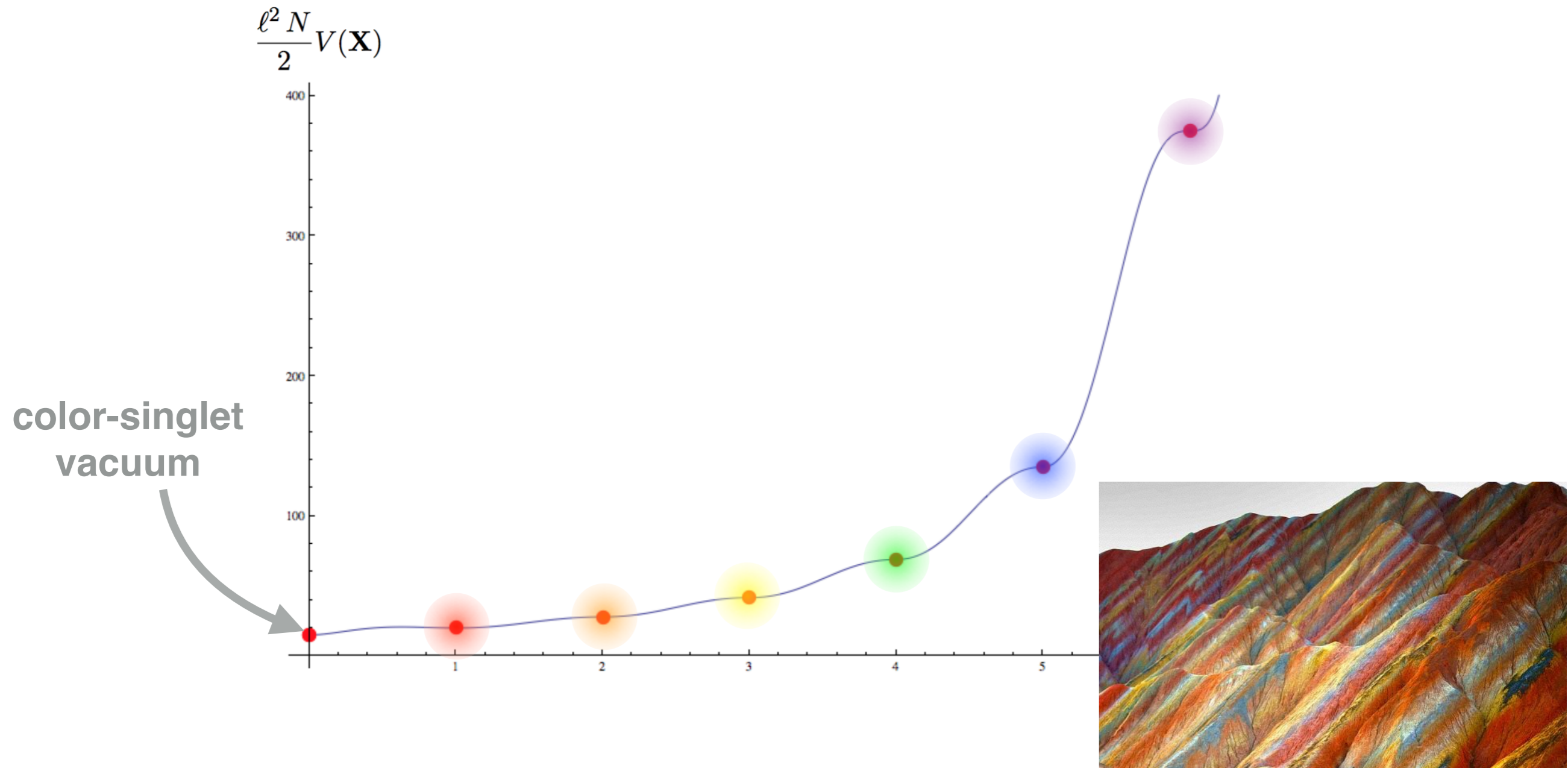
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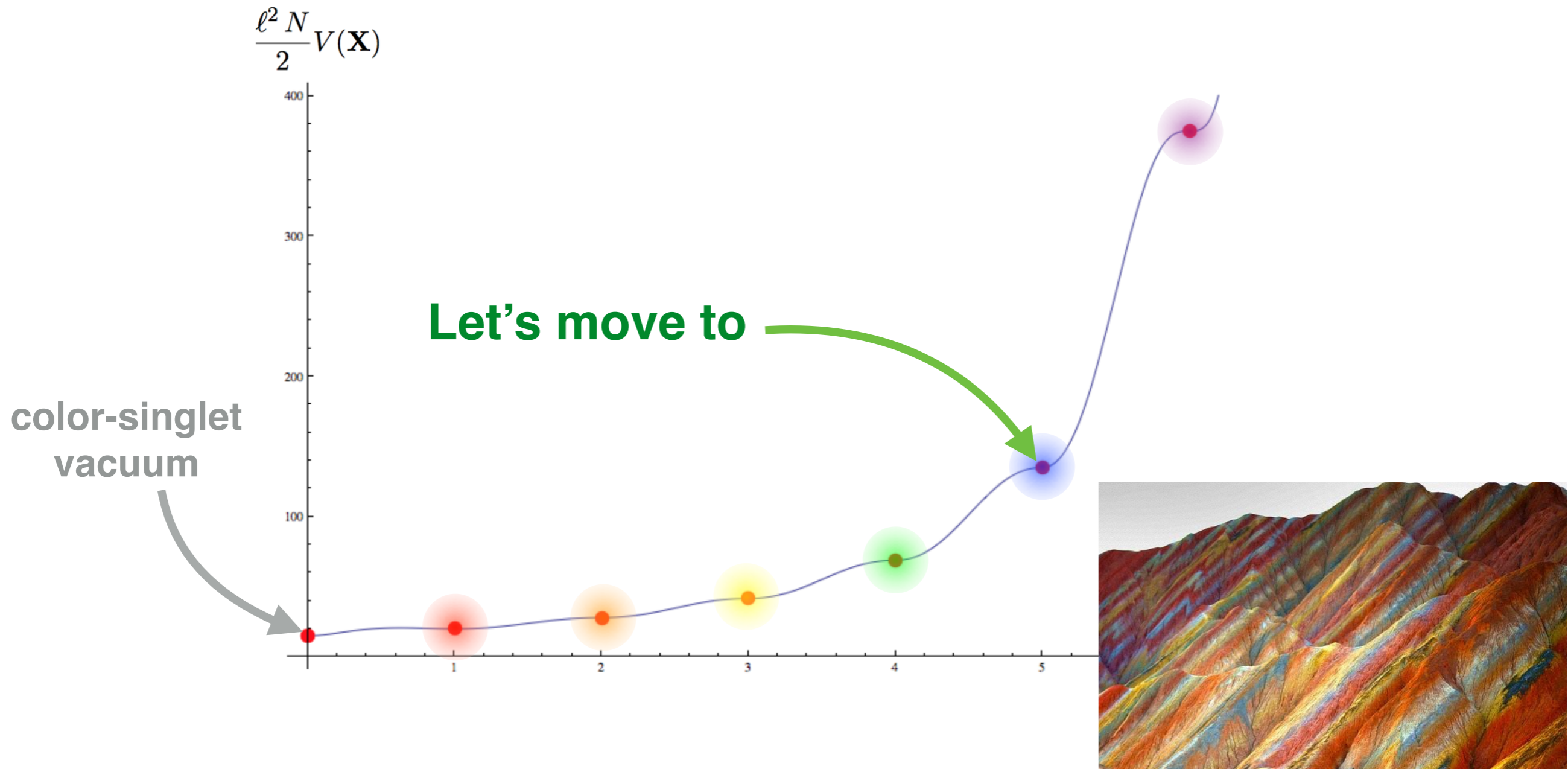
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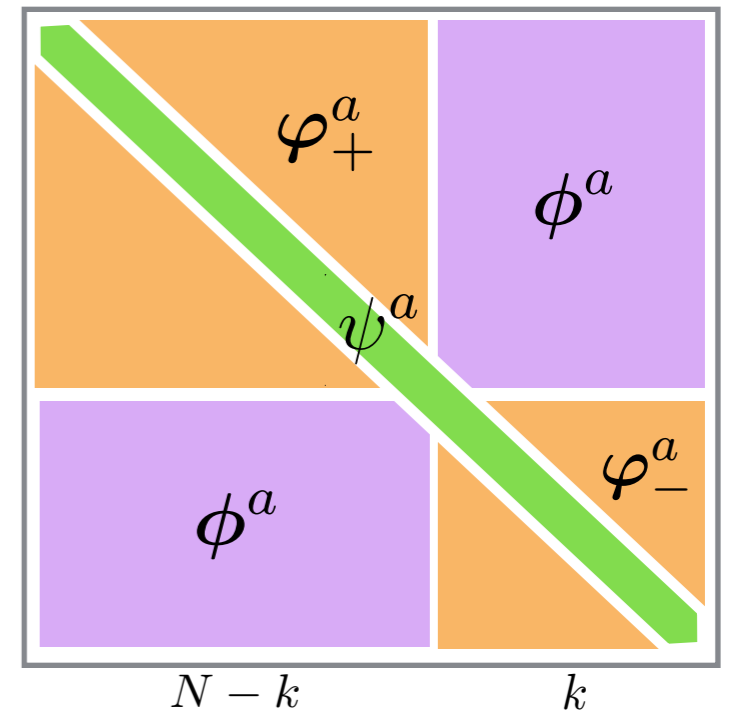


Symmetry Breaking



- Diagonal parts:

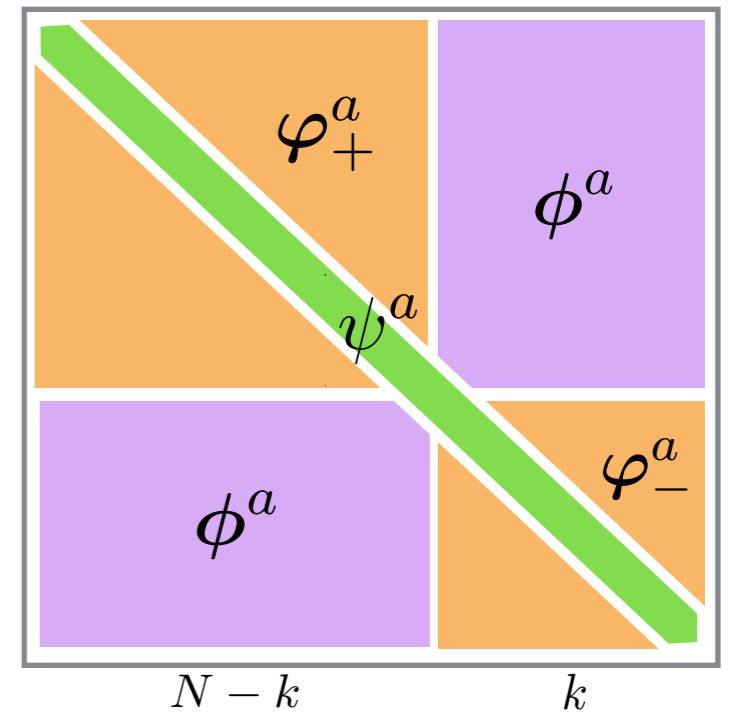


- Broken-sym. part:


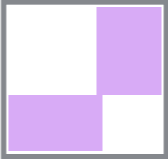


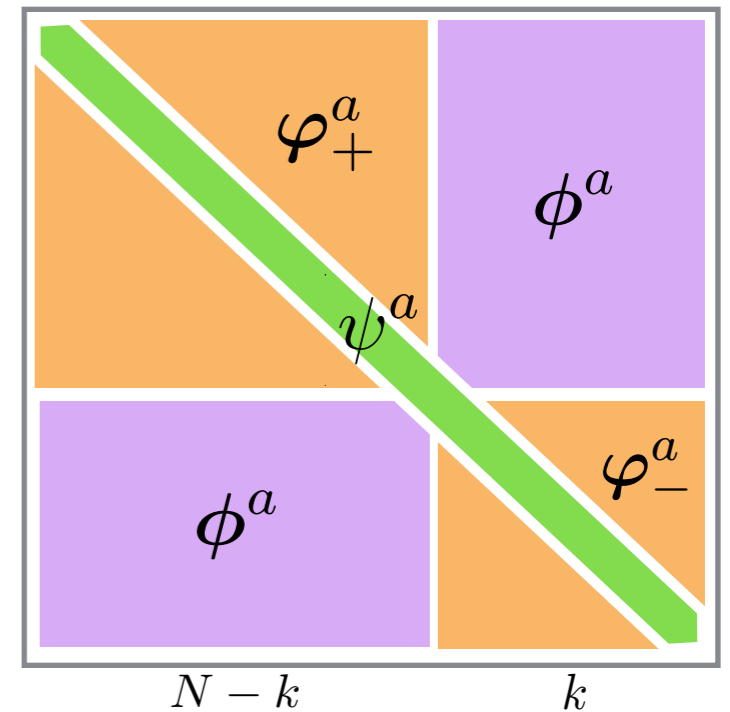
Symmetry Breaking

- Diagonal parts: 
 - ▶ **adjoint** in $SU(N - k) \times SU(k) \times U(1)$
 - ▶ still describe **massless** spin-two
- Broken-sym. part: 




Symmetry Breaking

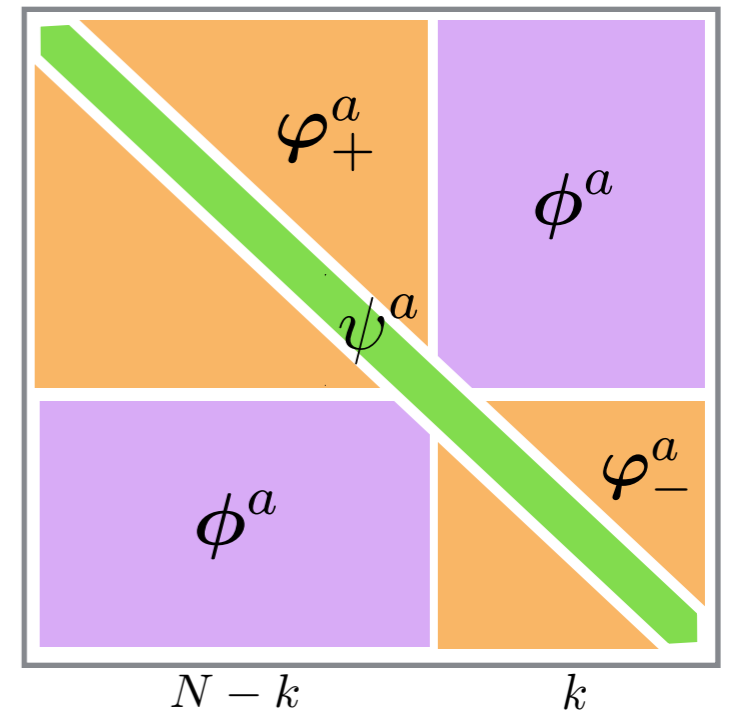
- Diagonal parts: 
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 - ▶ **bi-fundamental**



Symmetry Breaking

- Diagonal parts: 
 - ▶ adjoint in $SU(N - k) \times SU(k) \times U(1)$
 - ▶ still describe massless spin-two


- Broken-sym. part: 
 - ▶ bi-fundamental
 - ▶ combines with (or eaten by) spin-one field



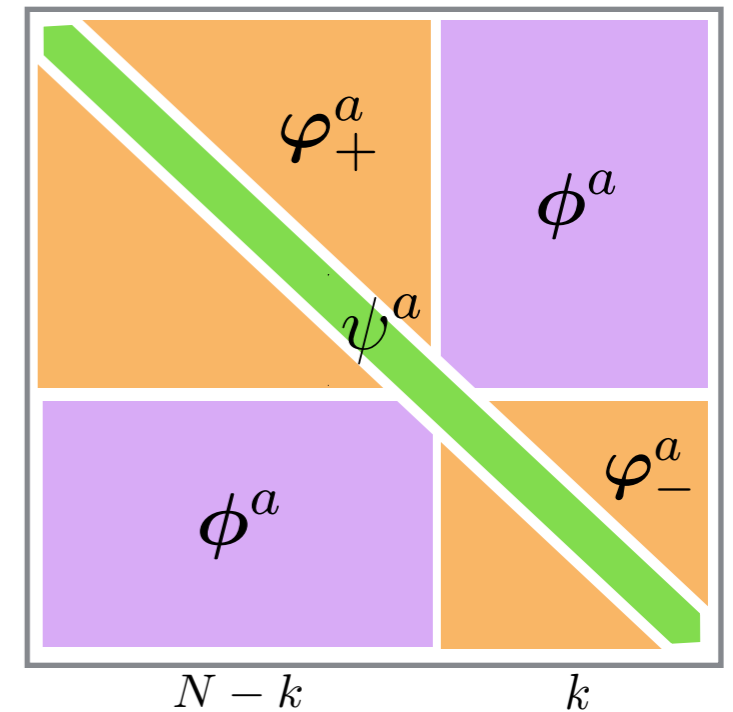
$$S_{\text{BS}}[\phi, \phi^a] = \int \phi \wedge \left(d\phi - \frac{1}{\ell} e_a \wedge \phi^a \right) - \phi_a \wedge \left(D\phi^a - \frac{1}{\ell} e^a \wedge \phi \right)$$

**Higgs-like
Mechanism!**

Symmetry Breaking

- Diagonal parts: 
 - ▶ **adjoint** in $SU(N - k) \times SU(k) \times U(1)$
 - ▶ still describe **massless** spin-two

- Broken-sym. part: 
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- ▶ combines with (or eaten by) spin-one field

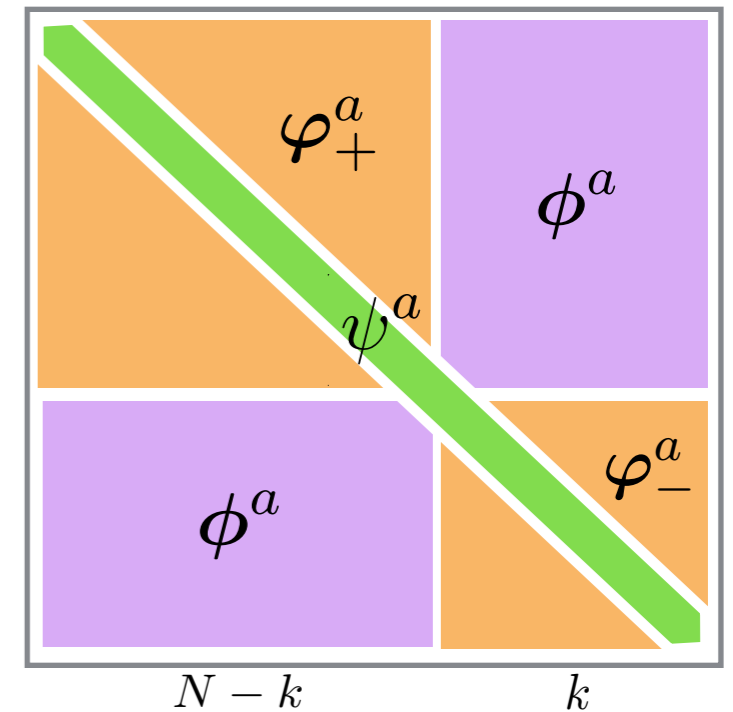
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**Higgs-like
Mechanism!**

- ▶ describe **partially-massless** spin-two

Symmetry Breaking

- Diagonal parts: 
 - ▶ **adjoint** in $SU(N - k) \times SU(k) \times U(1)$
 - ▶ still describe **massless** spin-two



- Broken-sym. part: 
 - ▶ **bi-fundamental**
 - ▶ combines with (or eaten by) spin-one field

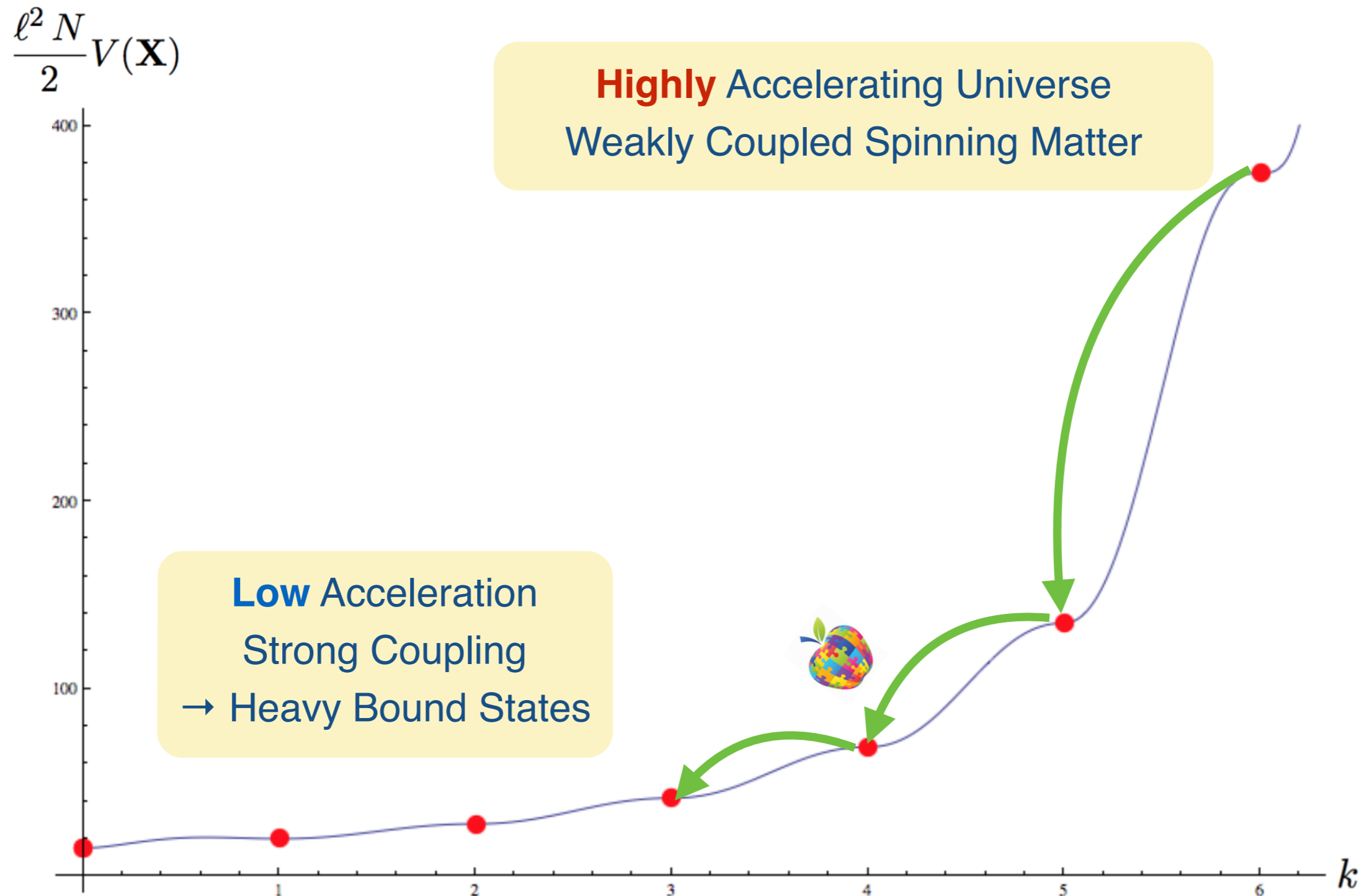
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**Higgs-like
Mechanism!**

- ▶ describe **partially-massless** spin-two

- **All Weakly Interacting for Large $k \sim N/2$**

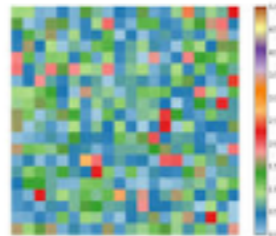
Cosmological Scenario



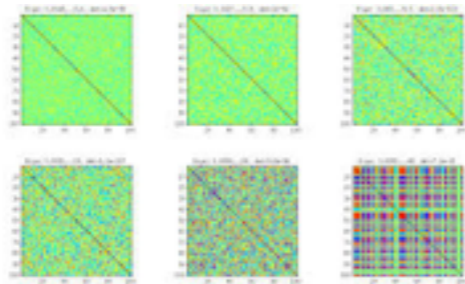
Quantum Colored Gravity

- Rainbow vacua contribution in the path integral:

Random Matrix Model



$$\mathcal{Z}_{\text{MM}} = \int d\mathbf{X} \exp [i c V(\mathbf{X})]$$



$$V(\mathbf{X}) = -\frac{2\sigma}{N\ell^2} \text{Tr} (\mathbf{I} + 3\mathbf{X}^2 + \mathbf{X}^3)$$


HS extension

- Color-decoration & Rainbow vacua extend to
 - ✓ 3D CS formulation of HS Gravity
 - ✓ Vasiliev Equations [to appear]

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- Color-decoration & Rainbow vacua extend to
 - ✓ 3D CS formulation of HS Gravity
 - ✓ Vasiliev Equations [to appear]
- Resulting spectrum after symmetry breaking
 - ▶ all the spins glue together to form an exotic one

Thank you

BRITISH LION  Presents

MONTE HALE
ADRIAN BOOTH
with THE SAGEBRUSH SERENADERS
and featuring OUTLAW - THE WILD HORSE

GRAVITY *from* **RAINBOW VALLEY**

TRADE SHOW:
AUG. 29
STUDIO ONE
11 a.m.

THE *First* OF A NEW SERIES
OF COLOUR WESTERNS
FROM REPUBLIC
IN
MAGNACOLOR
Greater THRILLS!
Greater ACTION!
Greater SPECTACLE!

A MAGNACOLOR PRODUCTION

The poster features a man in a cowboy hat and a woman standing next to a dark horse against a background of colorful, horizontal stripes. At the bottom, there is a scene of a stagecoach and riders in a desert landscape.